

TPC Gain Calibration and

Particle Identification

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Basics I: Bethe-Bloch

• particle with mass m and momentum p crosses a thin gas cell

specific energy loss due to ionisation → primary (cluster) charge

• mean:
$$\langle \frac{dE}{dx} \rangle = \frac{4\pi N e^4}{mc^2} \frac{z^2}{\beta^2} \Big(\frac{1}{2} \ln \frac{2mc^2 E_{max} \beta^2 \gamma^2}{I^2} - \frac{\beta^2}{2} - \frac{\delta(\beta)}{2} \Big)$$

Landau-like fluctuation:







Basics II: TPC clusters

- charges drift to the amplification region
- each pad row is considered as one gas cell





 \rightarrow 159 single dE/dx measurements (cluster charge Q)

Basics III: TPC signal and resolution

- the mean energy loss <dE/dx> is experimentally ill defined because of the tails in the energy loss distribution
 - \rightarrow average over the lowest m = η n values (truncated mean)

$$\langle S \rangle_{\eta} = \frac{1}{m} \sum_{i=0}^{m} Q_i$$

- \rightarrow it turns out empirically that <S> (TPC signal) follows a Gaussian distribution with the dE/dx resolution σ
- σ depends on the number of clusters n: $\sigma_{dE/dx}^2(n) = \sigma_{syst.}^2 + \frac{\sigma_{stat.}^2}{n}$.
- maybe the truncated mean will be replaced by a smooth weighting function

Gain calibration

- prerequisite for PID: the gain for each chamber and each pad must be equalized
- 3 different sources: Krypton, laser central electrode, pulser
- general tendencies within a chamber described by parabolic fits

	Amplitude	Fit	Convolution
Krypton	K _{amp}	K _{fit}	By default
Electrode	E _{amp}	E _{fit}	
Pulser	P _{amp}	P _{fit}	P _{conv}

 Krypton K_{amp} can do everything except: edge effect, chip-bychip fluctuations, regions which were not active during Kr runs

Edge effect

at the borders of a sector the gain is naturally lower

→ easiest solution: skip these pads, but we want to do better

Krypton cannot solve this problem, because of the 3cm range of the decay electrons



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General strategy for a gain map

- Krypton amplitude for one or several pads not o.k. → use fitted value
- remove chip-by-chip fluctuation by comparison of Krypton with artificially convoluted pulser data

$$A = K_{amp} \cdot \left(\frac{P_{amp}}{P_{fit}}\right)_{conv}$$

• remove the edge effect with the help of the central electrode

$$A = K_{fit} \cdot \frac{E_{amp}}{E_{fit}}$$



PID calibration



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The ALEPH parametrisation

Three theoretically possible ways to describe the mean TPC signal

• fit with the ALEPH parametrisation (our solution)

$$f(\beta\gamma) = \frac{P_1}{\beta^{P_4}} (P_2 - \beta^{P_4} - \ln(P_3 + \frac{1}{(\beta\gamma)^{P_5}}))$$

• fit with the Sternheimer parametrisation of the relativistic rise

$$-\langle \frac{dE}{dx} \rangle (\beta\gamma) = \frac{C_1}{\beta^2} \cdot \left(\ln(C_2\beta^2\gamma^2) - \beta^2 - \delta(\beta\gamma) \right) \qquad \delta = \begin{cases} 0, & X < X_0 \\ 4.6052X + C_0 + a(X_1 - X)^m, & X_0 < X < X_1 \\ 4.6052X + C_0, & X > X_1 \end{cases}$$

 don't fit at all, learn sth. about the energy loss of charged particles and use a tuned physics model
 → see talks of P. Christiansen

PID calibration with cosmics data (1)

5,000,000 cosmic events from June 2008, simple Kr calibration



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PID calibration with cosmics data (2)

- The lines shown here are all from the ALEPH parametrisation with the same parameter set, only the rest mass of the particle is input
- The fitting procedure is work in progress:

→ start from the Aleph parameters and hand-adjust them slightly

 \rightarrow use this basic PID to plot it is a function of $\beta\gamma$ and refit

• $1/\beta^2$ part seems to be slightly steeper compared to ALEPH and $I_{plat}/I_0 \approx 1.51$



P1 = 3.61; P2 = 16.15; P3 = 1.32e-7; P4 = 2.49; P5 = 2.99

dE/dx resolution σ with cosmics data

- cosmic tracks are reconstructed twice (upper and lower half of the TPC)
- by comparing their dE/dx, we can estimate the resolution (see Marian's talk in the physics forum last week)



 Please note: angular corrections and for the z-dependence are still done on the basis of MC data; results also hopefully very soon

Dependence of σ on n in MC data

 pions in the MIP region are well separated from other species, so this dependence can be fixed there



currently only MC data, but hopefully very soon with cosmcis

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Conclusions

- cosmic data already gives us a very good idea of how the beam data will look like
- a fit with the ALEPH parametrisation seems to be very well suited to describe the energy loss of all particle species
- we can extract a first set parameters which we can then further improve as soon as beam-gas data is available
- we now know which type of particles we can find in our cosmic events...



Backup slides









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