



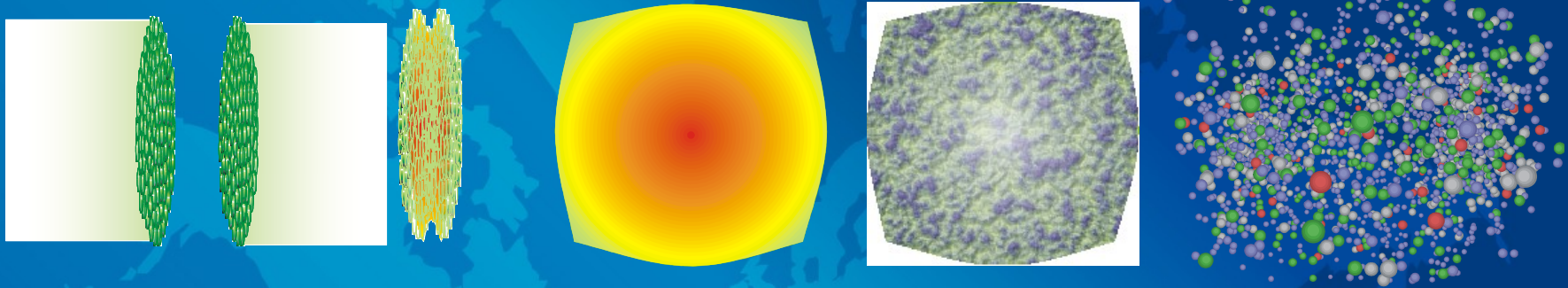
Space-time evolution of the Quark Gluon Plasma

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High-energy nucleus-nucleus Collisions



High-Energy Nuclear Collisions



Time →

Plot: Steffen A. Bass, Duke University

1) Initial condition:

- Baryon transfer
- E_T production
- Partonic dof

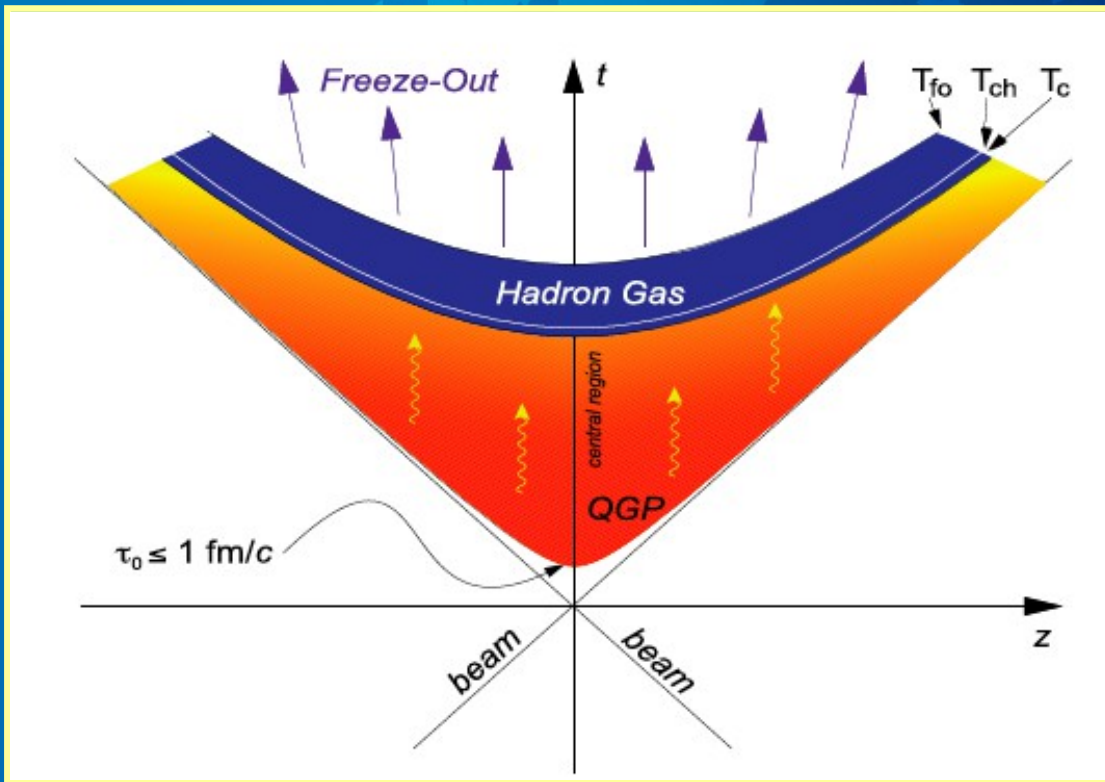
2) System evolves:

- parton/hadron expansion

3) Bulk freeze-out

- hadronic dof
- **inel. interactions cease:**
particle ratios, T_{ch}, μ_B
- **elas. interactions cease**
Particle spectra, $T_{th}, \langle \beta_T \rangle$

Space-time evolution



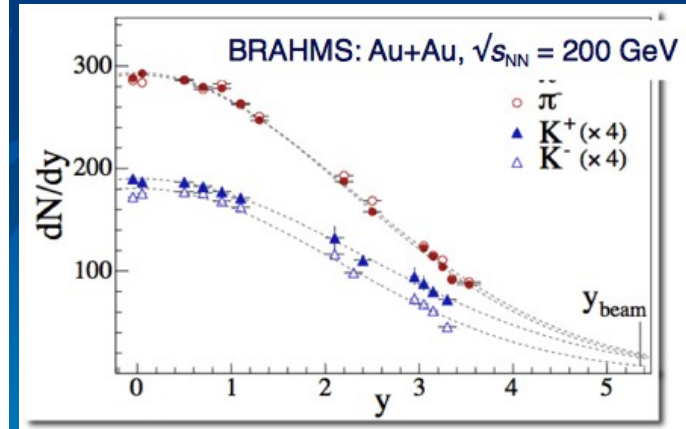
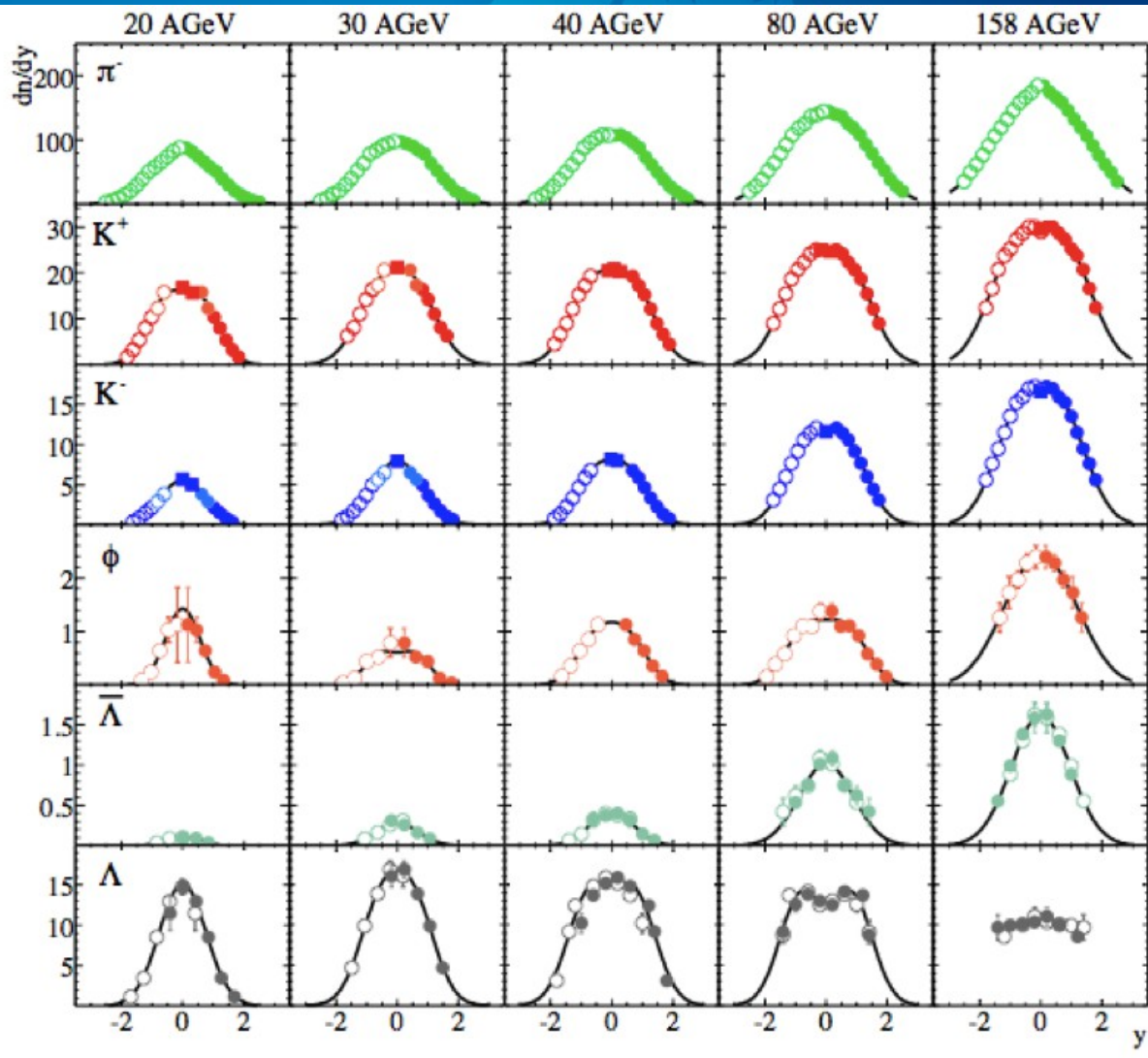
Plot: courtesy of R. Stock.

- **QGP life time**
 $10 \text{ fm}/c \approx 3 \cdot 10^{-23} \text{ s}$
- **thermalization time**
 $0.2 \text{ fm}/c \approx 7 \cdot 10^{-25} \text{ s}$
→ hydrodynamical expansion until freeze-out
simplest model: only longitudinal expansion, 1d
→ Bjorken model
- **collision time**
 $2R/\gamma = 0.005 \text{ fm}/c$
 $\approx 2 \cdot 10^{-26} \text{ s}$

Outline

- Introduction
- Longitudinal expansion – Bjorken picture
- Transverse expansion
 - transverse radial flow
 - transverse elliptic flow v_2
 - higher harmonics, v_3, v_4, v_5, \dots
- Hydrodynamical model description
- Summary

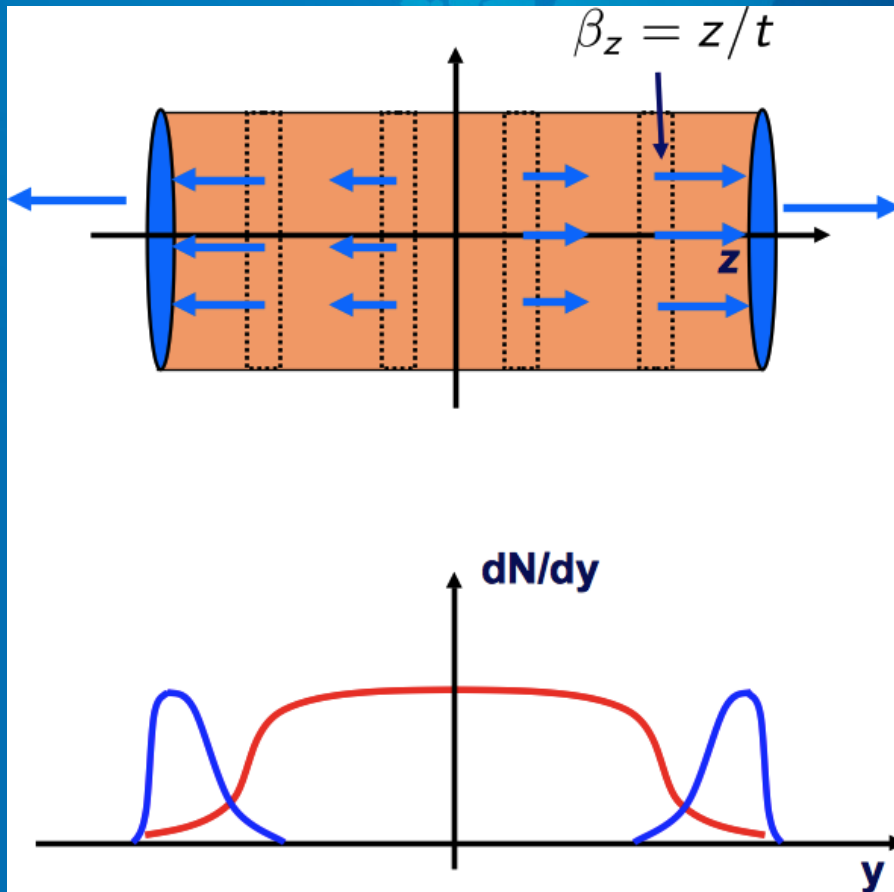
Rapidity distribution in A-A



With increasing collision energy:

- wider distribution,
- becomes flatter around mid-rapidity

Bjorken model



Velocity of the local system
at position z at time t :

$$\beta_z = z/t$$

Proper time τ in this system:

$$\begin{aligned}\tau &= t/\gamma = t\sqrt{1 - \beta^2} \\ &= \sqrt{t^2 - z^2}\end{aligned}$$

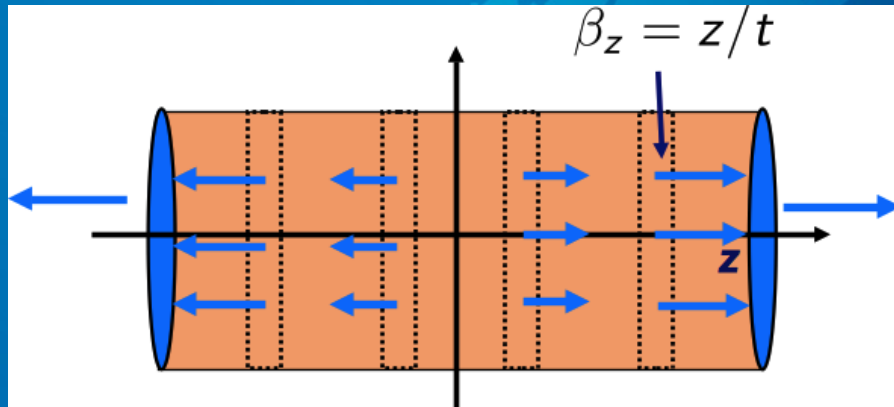
In the Bjorken model all
thermodynamic quantities only
depend on τ , e.g., the particle
density:

$$n(t, z) = n(\tau)$$

This leads to a constant rapidity
density of the produced particles
(at least at central rapidities):

$$\frac{dN_{ch}}{dy} = \text{const.}$$

1d - Bjorken model (I)



The 1D Bjorken model is based on the assumption that dN_{ch}/dy is constant (around mid-rapidity). This means that the central region is invariant under Lorentz transformation. This implies $\beta_z = z/t$ and that all thermodynamic quantities depend only on the proper time τ

Initial conditions in the Bjorken model:

$$\varepsilon(\tau_0) = \varepsilon_0, \quad u^\mu = \frac{1}{\tau_0} (t, 0, 0, z) = \frac{x^\mu}{\tau_0}$$

Initial energy density

In this case the equations of ideal hydrodynamics simplify to

$$\frac{d\varepsilon}{d\tau} + \frac{\varepsilon + p}{\tau} = 0$$

$\varepsilon = E/V$: energy density
p	: pressure
$s = S/V$: entropy density

Bjorken model (II)

For an ideal gas of quarks and gluons, i.e., for

$$\varepsilon = 3p, \quad \varepsilon \propto T^4$$

This leads to

$$\varepsilon(\tau) = \varepsilon_0 \left(\frac{\tau}{\tau_0} \right)^{-4/3}, \quad T(\tau) = T_0 \left(\frac{\tau}{\tau_0} \right)^{-1/3}$$

The temperature drops to the critical temperature at the proper time

$$\tau_c = \tau_0 \left(\frac{T_0}{T_c} \right)^3$$

And thus the lifetime of the QGP in the Bjorken model is

$$\Delta\tau_{\text{QGP}} = \tau_c - \tau_0 = \tau_0 \left[\left(\frac{T_0}{T_c} \right)^3 - 1 \right]$$

QGP lifetime in 1d - Bjorken model

$$\varepsilon_0 = 11 \text{ GeV}/\text{fm}^3 = 11 \cdot 0.197^3 \text{ GeV}^4 \quad \text{for } \tau_0 = 1\text{fm}/c$$

$$1 = \hbar c = 0.197 \text{ GeV} \cdot \text{fm}$$

$$\varepsilon_0 = g_{\text{QGP}} \frac{\pi^2}{30} T^4 \quad \rightarrow \quad T_0 = \left(\frac{30 \varepsilon}{\pi^2 g} \right)^{1/4}$$

Parameters

$$\varepsilon_0 \tau = 3 \text{ GeV}/\text{fm}^2$$

ΔT_{QGP}

$$0.84 \text{ fm}/c$$

$$\varepsilon_0 \tau = 5 \text{ GeV}/\text{fm}^2$$

$$1.70 \text{ fm}/c$$

$$\varepsilon_0 \tau = 11 \text{ GeV}/\text{fm}^2$$

$$3.9 \text{ fm}/c$$

Fixed parameters: $N_f = 2$, $T_c = 170 \text{ MeV}$, $\tau_0 = 1 \text{ fm}/c$

Quick estimate for LHC

Bjorken formula: $\varepsilon \cdot \tau_0 = \frac{\langle m_T \rangle}{A} \frac{dN}{dy} \Big|_{y=0}$

Transverse area in collisions with $b \approx 0$:

$$A \approx \pi R_{\text{Pb}}^2 = \pi (6.62 \text{ fm})^2 \approx 140 \text{ fm}^2$$

Estimate for the mean transverse momentum:

$$\langle p_T \rangle = 0.66 \text{ GeV}/c \rightsquigarrow \langle m_T \rangle \approx \sqrt{(0.138 \text{ GeV})^2 + (0.66 \text{ GeV})^2} = 0.67 \text{ GeV}$$

Measured charged particle multiplicity:

$$dN_{ch}/d\eta \approx 1601 \pm 60 \quad (5\% \text{ most central})$$

$$\rightsquigarrow \frac{dN}{dy} \Big|_{y=0} = \frac{3}{2} \cdot \underbrace{\left(1 - \frac{m^2}{\langle m_T \rangle^2}\right)^{-1/2}}_{1.02} \cdot \frac{dN_{ch}}{d\eta} \Big|_{\eta=0} = 2450 \pm 92$$

1.02 Larger (up to ≈ 1.2) if p and K are taken into account

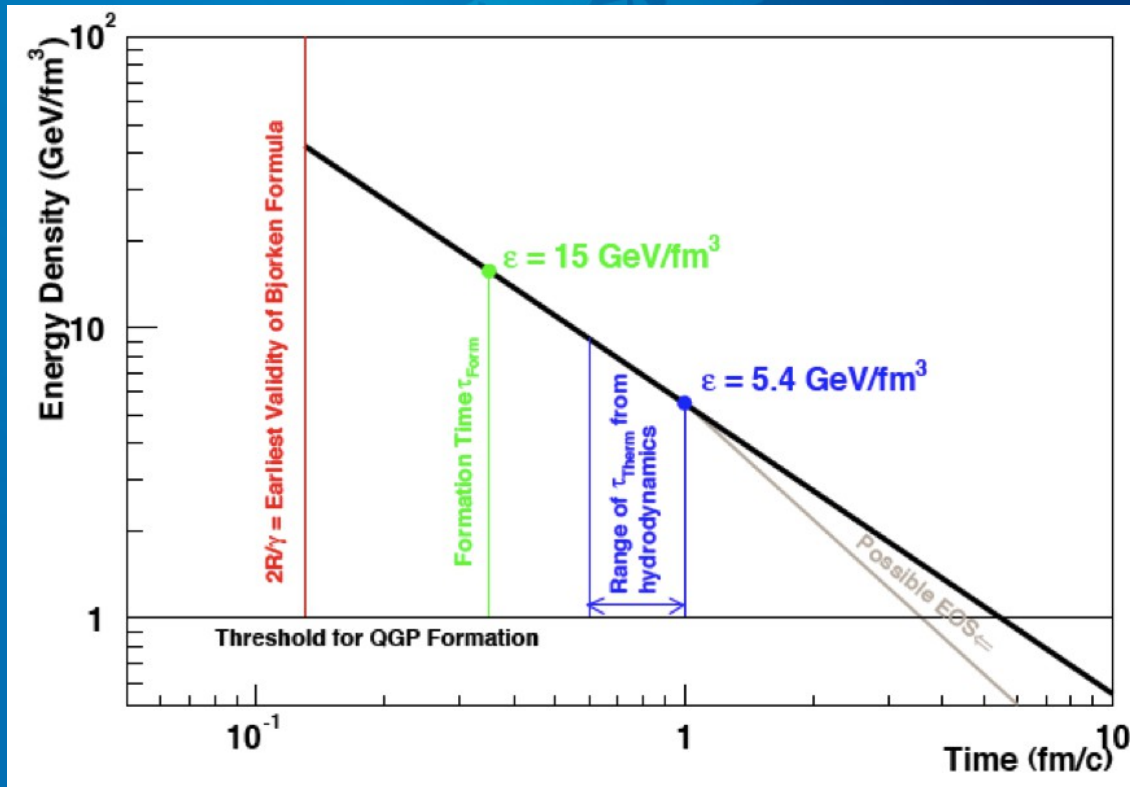
LHC:

$$\varepsilon \cdot \tau_0 = (11.7 \pm 0.43) \text{ GeV}/\text{fm}^2 \quad (\text{Pb+Pb}@ \sqrt{s_{NN}} = 2.76 \text{ TeV})$$

RHIC:

$$\varepsilon \cdot \tau_0 \approx 5 \text{ GeV}/\text{fm}^2 \quad (\text{Au+Au}@ \sqrt{s_{NN}} = 0.2 \text{ TeV})$$

Energy density evolution in 1d-Bjorken



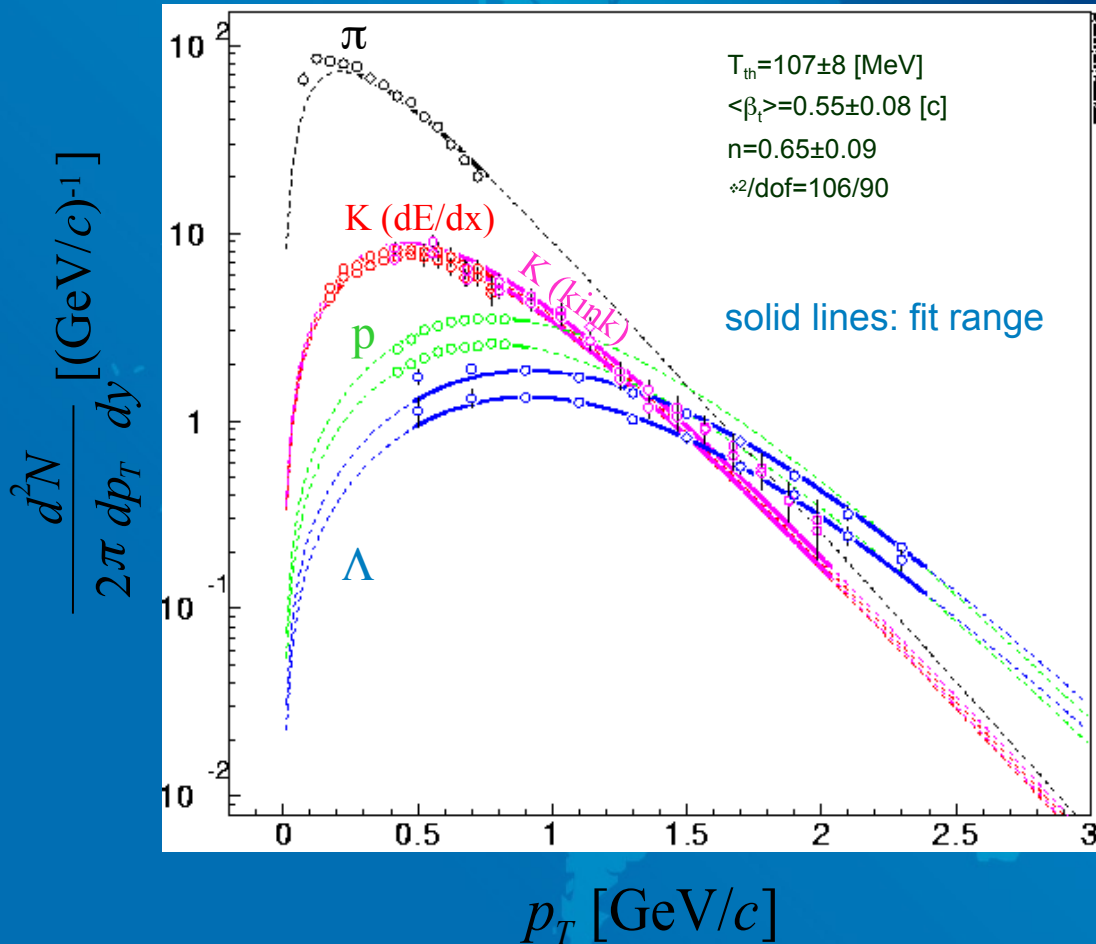
$\tau_0 = 1 \text{ fm}/c$ is generally considered as a conservative estimate for the use in the Bjorken formula.
Other estimates yields shorter times (e.g. $\tau_0 = 0.35 \text{ fm}/c$) resulting in initial energy densities at RHIC of up to $15 \text{ GeV}/\text{fm}^3$

A faint, light blue world map is centered on the slide, showing the outlines of continents and oceans. The map is semi-transparent, allowing the text to be clearly visible over it.

Transverse Expansion

Transverse radial flow: particle spectra

Particle Spectra*



*Au+Au @130 GeV, STAR

- Typical **mass ordering** in inverse slope from **light** π to **heavier** Λ

- Two-parameter fit describes yields of π, K, p, Λ

- $T_{th} = 90 \pm 10 \text{ MeV}$
- $\langle \beta_t \rangle = 0.55 \pm 0.08 c$

⇒ Disentangle

collective motion from thermal random walk

Thermal Model + Radial Flow Fit

Source: each volume element is assumed to be

– in local **thermal equilibrium**: T_{fo}

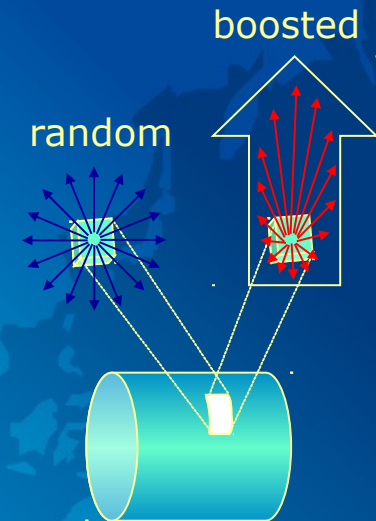
– **boosted** in transverse radial direction: $\rho = f(\beta_s)$

E.Schnedermann, J.Sollfrank, and U.Heinz, Phys. Rev. C48, 2462(1993).

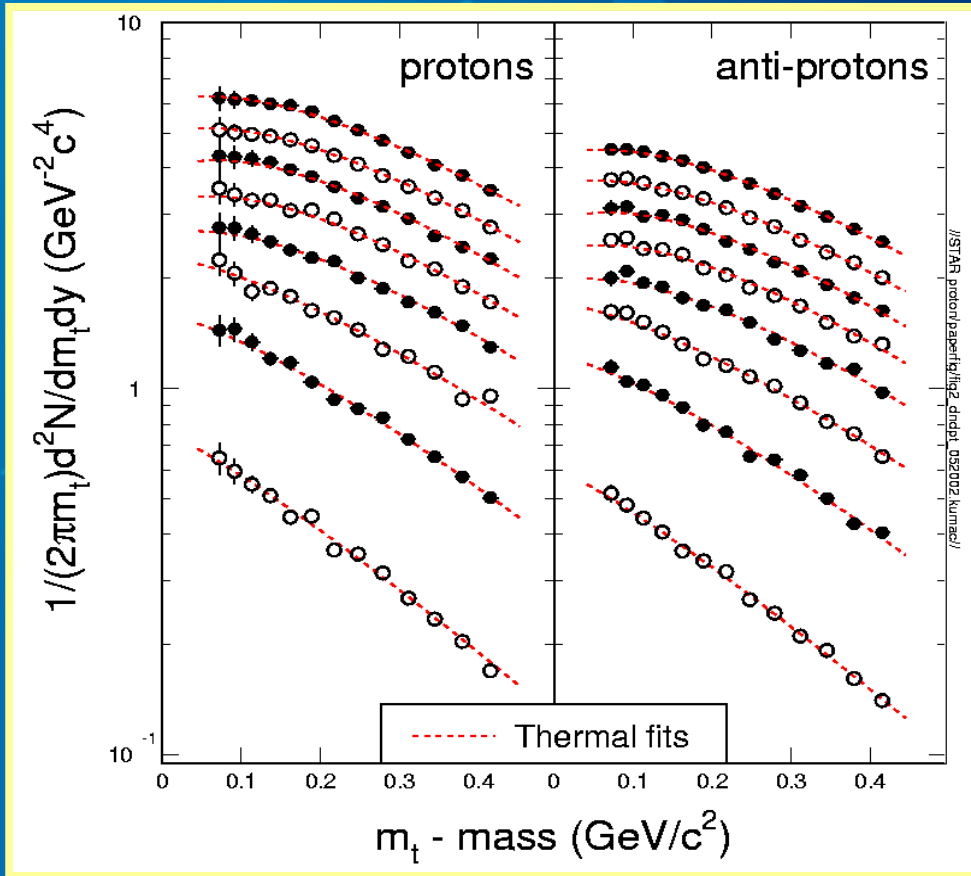
$$E \frac{d^3 N}{dp^3} \propto \int_{\sigma} e^{-(u^\mu p_\mu)/T_{fo}} p d\sigma_\mu \Rightarrow$$

$$\frac{dN}{m_T dm_T} \propto \int_0^R r dr m_T K_1 \left(\frac{m_T \cosh \rho}{T_{fo}} \right) I_0 \left(\frac{p_T \sinh \rho}{T_{fo}} \right)$$

$$\rho = \tanh^{-1} \beta_T \quad \beta_T = \beta_s \left(\frac{r}{R} \right)^\alpha \quad \alpha = 0.5, 1, 2$$



(anti-)Protons From RHIC Au+Au@130GeV



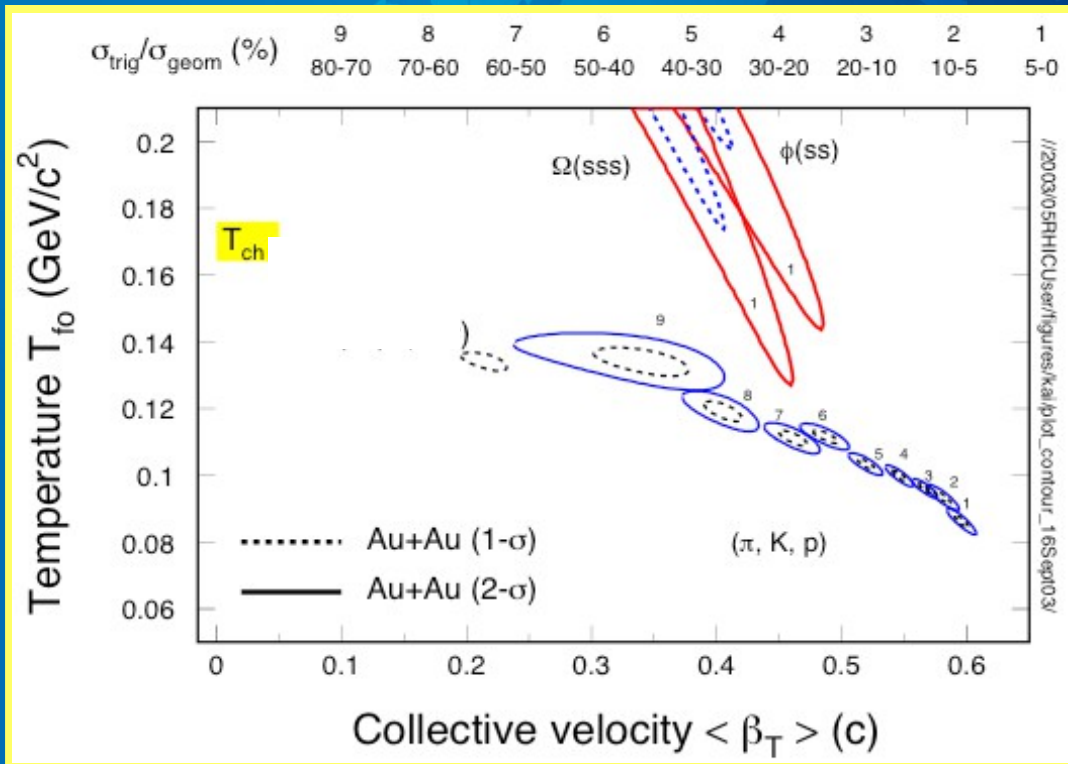
$$m_T = \sqrt{p_T^2 + \text{mass}^2}$$

Centrality dependence:

- **spectra** at low momentum de-populated, become **flatter** at larger momentum

ü stronger **collective flow** in more central collisions, $\langle \omega_t \rangle = 0.55 \pm 0.08$

Kinetic Freeze-out at RHIC



STAR Data: Nucl. Phys. A757, (2005 102),

*A. Baran, W. Broniowski and W. Florkowski, Acta. Phys. Polon. B 35 (2004) 779.

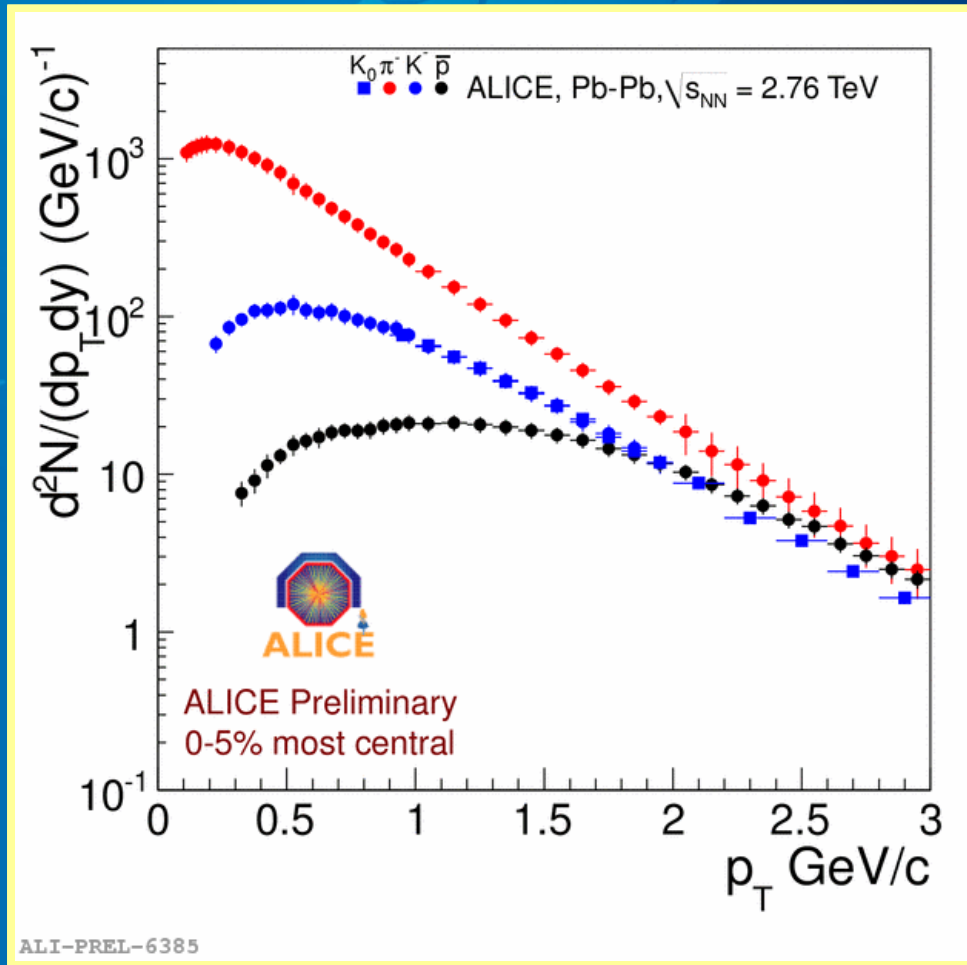
0) T and β_T are anti-correlated

1) Multi-**strange hadrons** ϕ and Ω **freeze-out earlier** than (π, K, p)
 ü Collectivity prior to hadronization

2) Sudden single freeze-out*:
 Resonance decays lower T_{fo} for (π, K, p)
 ü **Collectivity prior to hadronization**

ü **Partonic Collectivity ?**

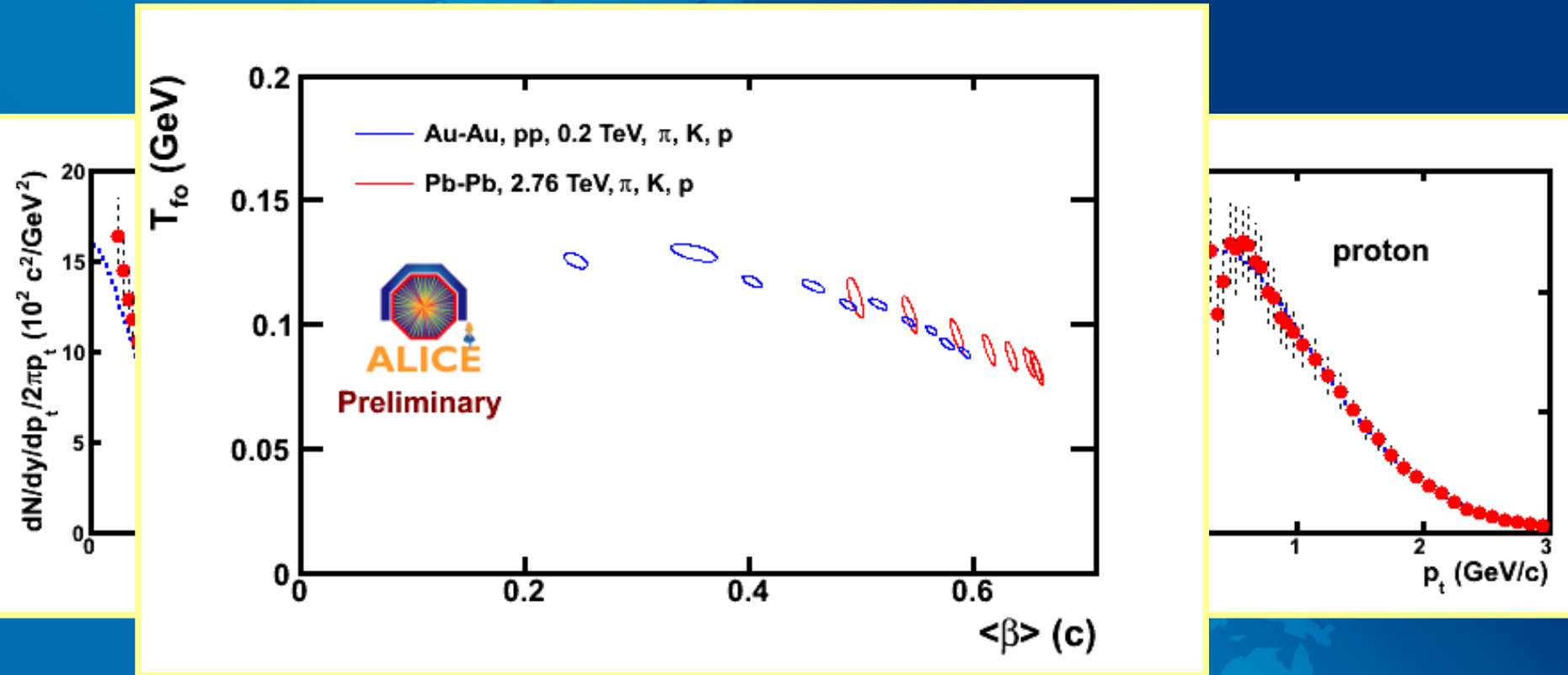
LHC: Identified particle spectra



Spectra harder at LHC

\Rightarrow **stronger** collective
flow at LHC than at
RHIC

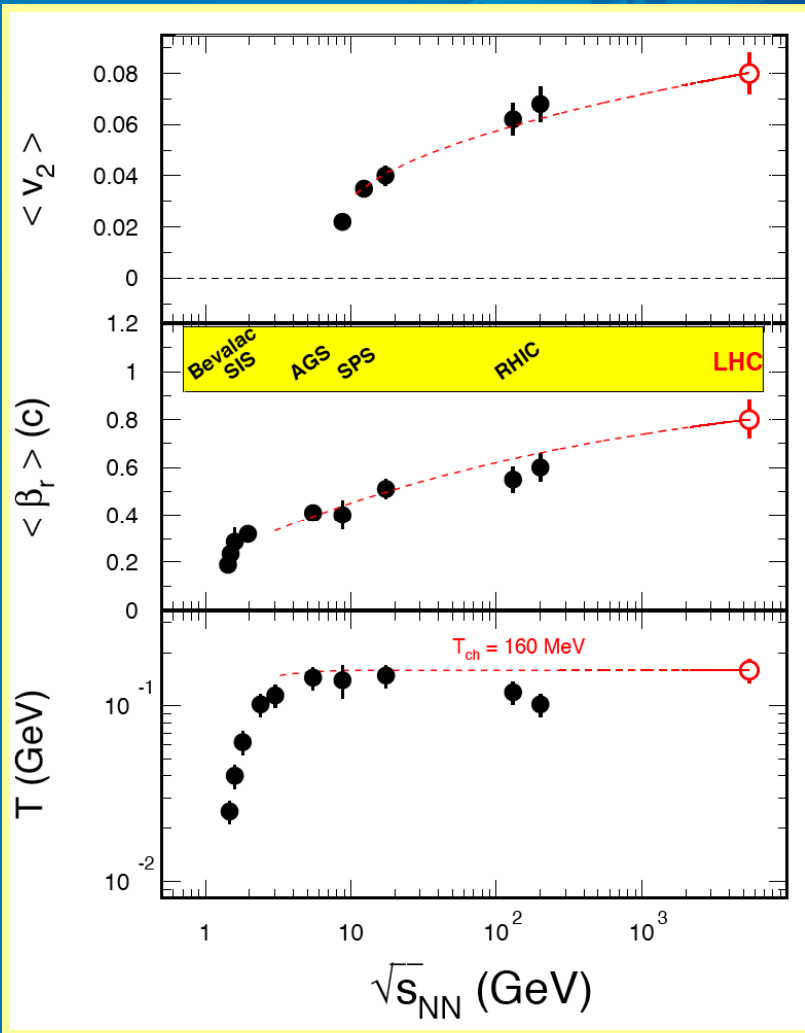
Collective expansion



Blast wave parametrization **describes spectra** at **10% level**

Collective **flow velocity increases** from RHIC to LHC by **10%**

Collective Flow - Energy Dependence



Collectivity parameters $\langle \beta_T \rangle$

and $\langle v_2 \rangle$ **increase** with

collision energy

strong collective expansion

at RHIC !

$\langle \beta_T \rangle_{\text{RHIC}} \gg 0.6$

expected **strong partonic**

expansion at LHC,

$\langle \beta_T \rangle_{\text{LHC}} \approx 0.8, T_{fo} \approx T_{ch}$

Lesson I

- At LHC, Initial energy density approx. $50 \text{ GeV}/\text{fm}^3$
- much larger than critical energy density $\varepsilon_c = 0.7 \text{ GeV}/\text{fm}^3$
- Strong **collective expansion**, $\langle \beta_T \rangle = 0.6 - 0.7$
at **highest collider energies**
- Particles carrying **strange quarks** show that **collective expansion** develops **before hadronization**, $\langle \beta_T \rangle = 0.3 - 0.4$
– among quarks and gluons (?)



Transverse Expansion

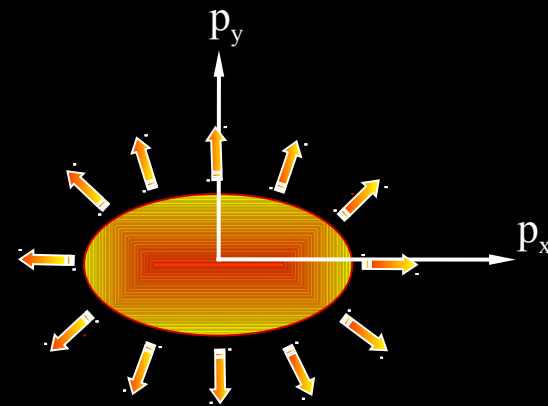
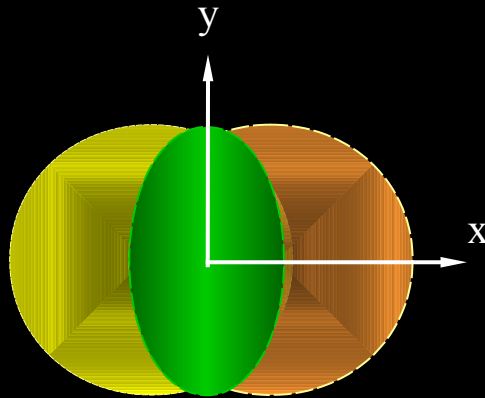
Transverse elliptic flow: event anisotropy

Anisotropy Parameter v_2

coordinate-space-anisotropy

ü

momentum-space-anisotropy

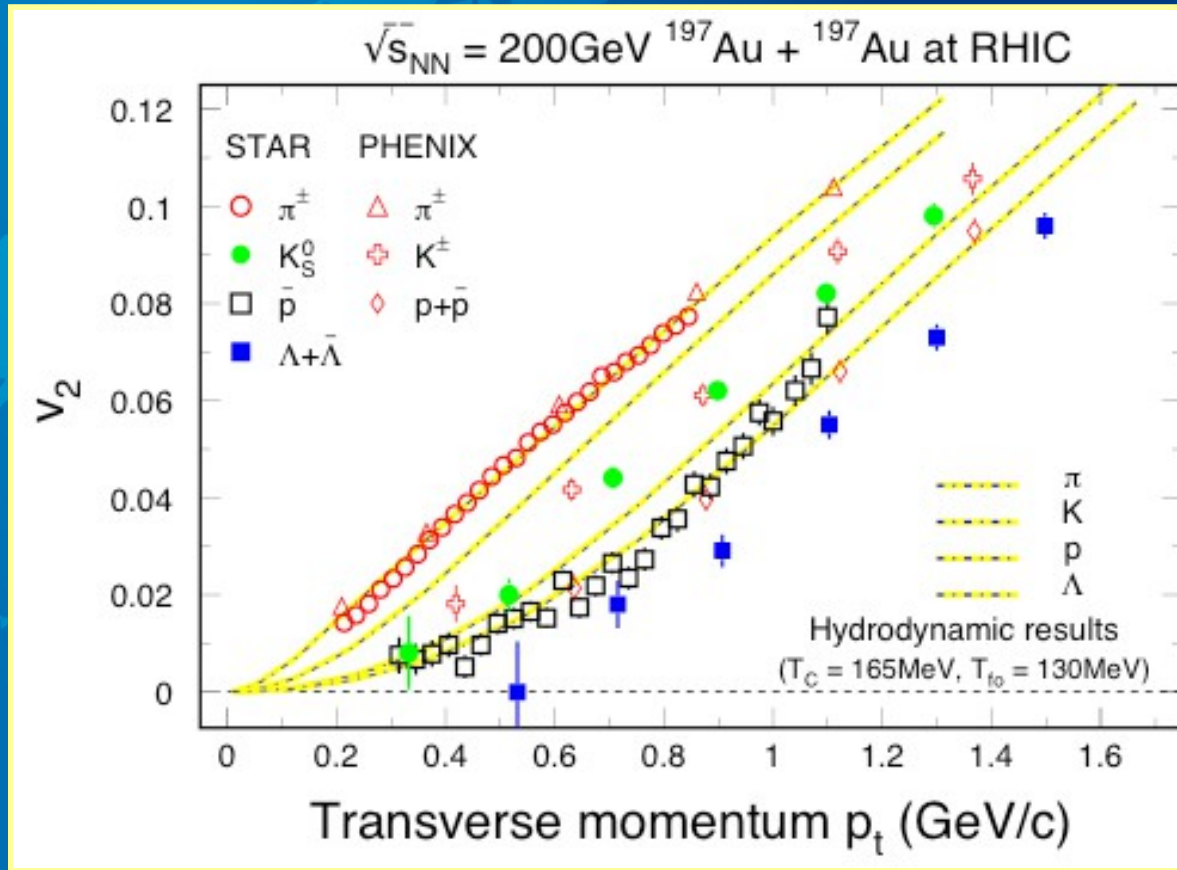


$$\varepsilon = \frac{\langle y^2 - x^2 \rangle}{\langle y^2 + x^2 \rangle}$$

$$v_2 = \langle \cos 2\varphi \rangle, \quad \varphi = \tan^{-1}\left(\frac{p_y}{p_x}\right)$$

Initial/final conditions, EoS, degrees of freedom

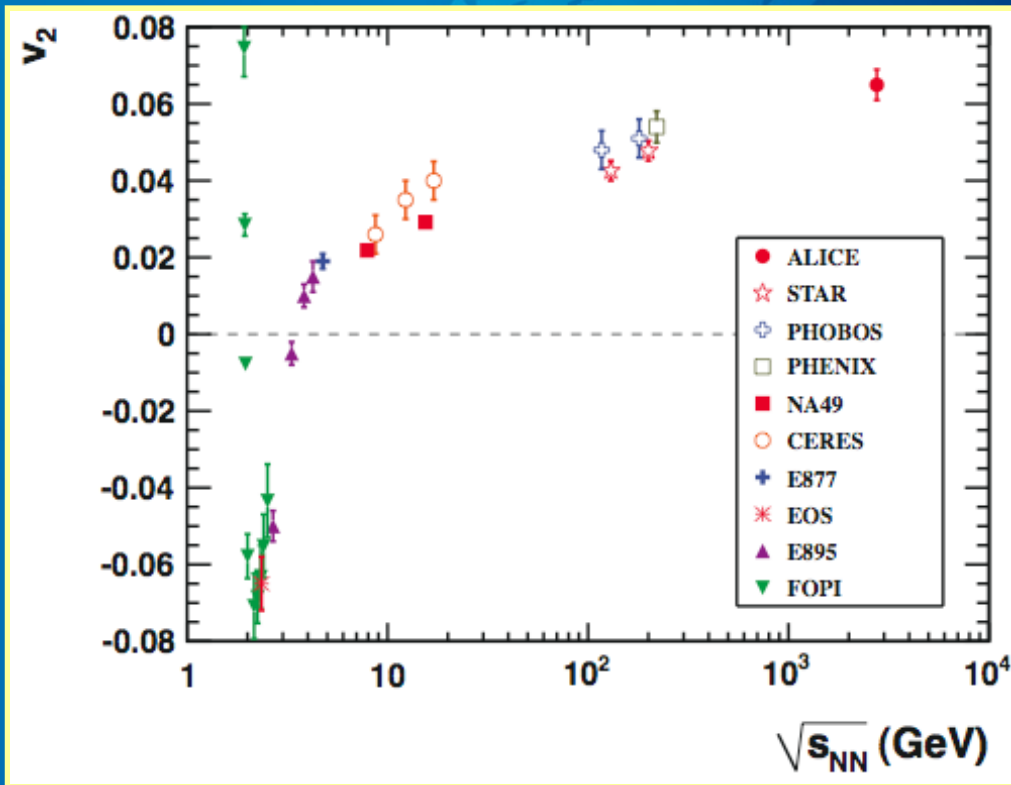
v_2 in the low- p_T Region



P. Huovinen, private communications, 2004

- v_2 approx. linear in p_T , mass ordering from light π to heavier Λ
- characteristic of hydrodynamic flow, sensitive to EOS !

Elliptic flow in ALICE



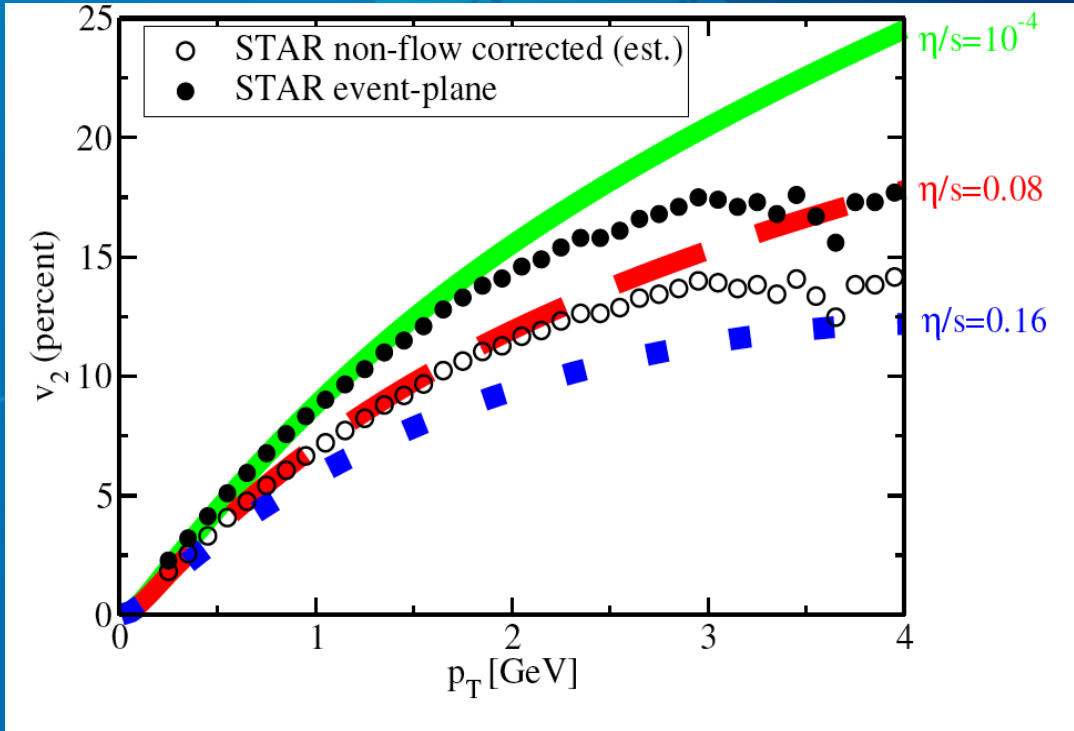
ALICE, submitted for publication, arXiv:1011.3914 [nucl-ex].

$\sqrt{s_{NN}} > \sim 4$ GeV: initial
eccentricity leads to pressure
gradients that cause positive v_2

$2 < \sqrt{s_{NN}} < 4$ GeV: velocity of
the nuclei is small so that
presence of spectator matter
inhibits in-plane particle
emission ("squeeze-out")

$\sqrt{s_{NN}} < 2$ GeV: rotation of the
collision system leads to
fragments being emitted in-
plane

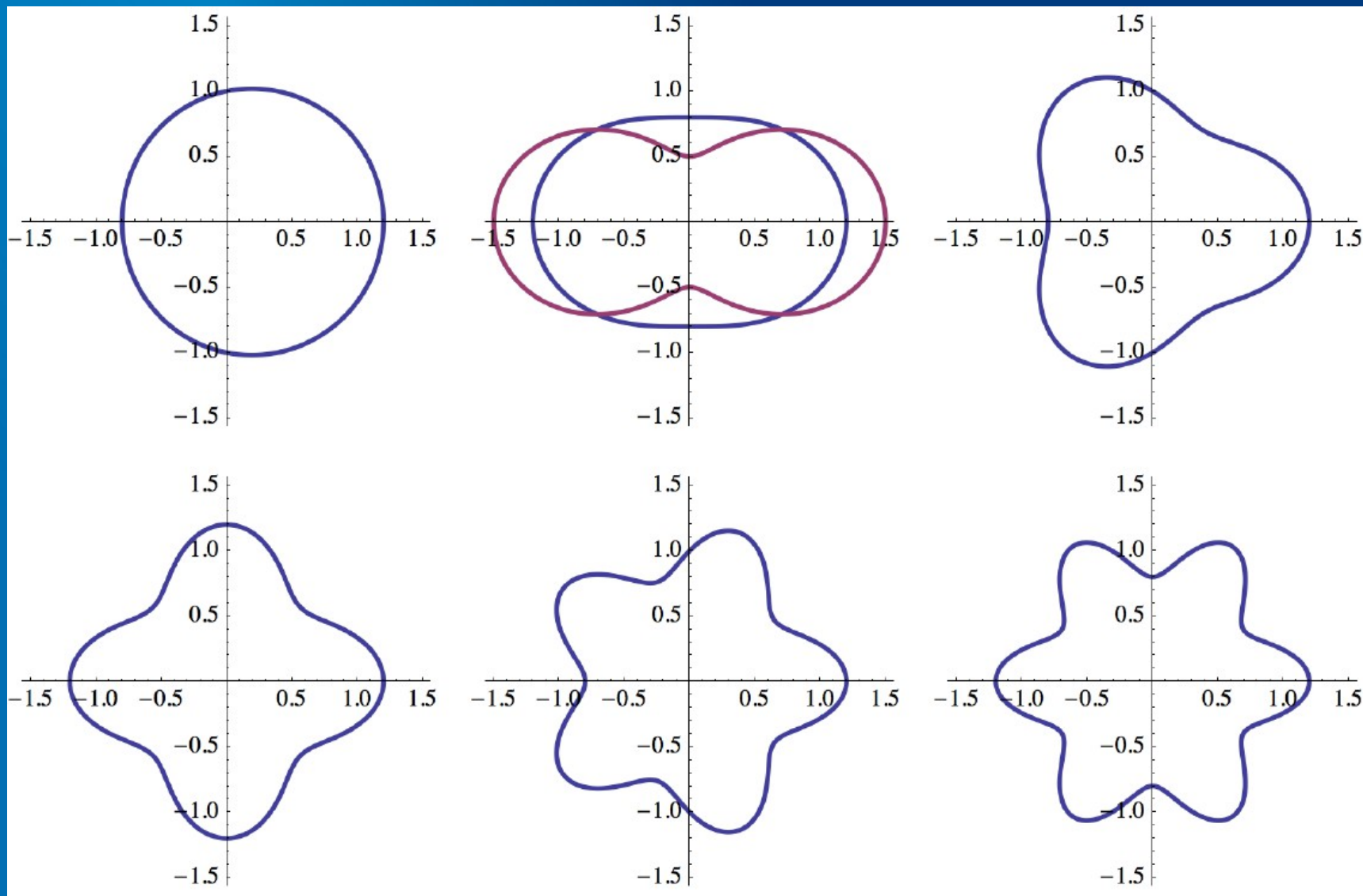
Non-ideal Hydro-dynamics



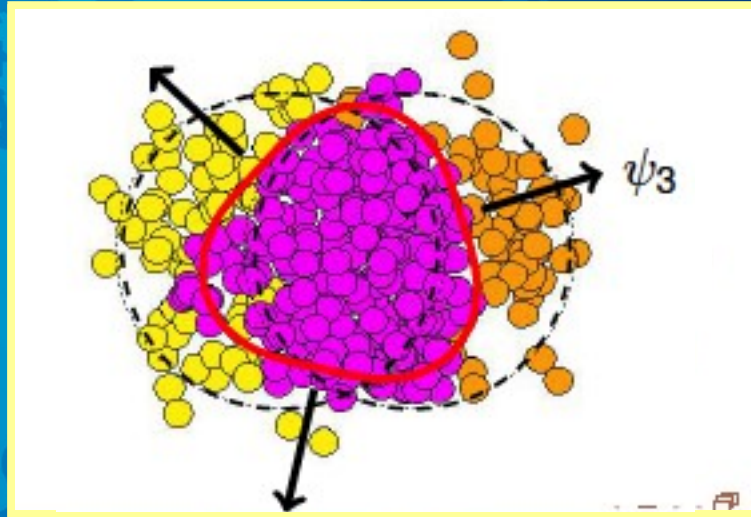
M.Luzum and R. Romatschke, PRC 78 034915 (2008); P. Romatschke, arXiv:0902.3663.

- Spectra and flow reproduced by ideal hydrodynamics calcs.
- Shear viscosity to entropy density ratio close to AdS/CFT bound
- viscosity leads to decrease in v_2 , ultra-low viscosity sufficient to describe data
- Hydro-limit exceeded at LHC ?

$$f(\varphi) = 1 + 2v_n \cos(n\varphi)$$



Can there be v_3 ?



figs.: courtesy of M. Luzum.

- reaction plane \perp participant plane
- fluctuating initial state is seed for v_3
- can CGC be challenged ?

higher harmonics

- extract power spectrum of v_n , like Planck*
- higher harmonics
- odd harmonics important
- v_3 : access η/s
- Higher harmonics strongly damped ($v_n, n>10 = 0$)

STAR, arXiv:1301.2187 [nucl-ex];

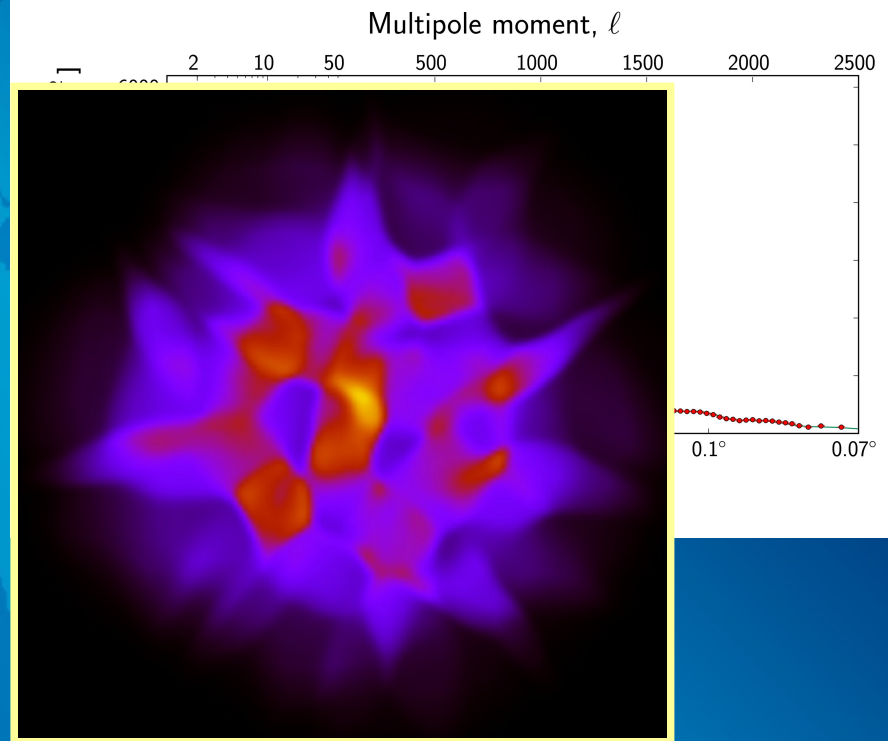
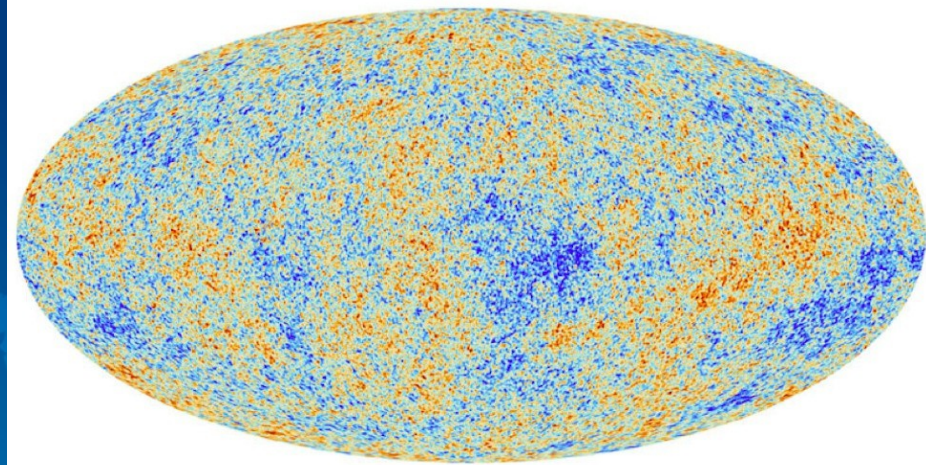
STAR, PRL 92 (2004) 062301;

A.Mocsy and P. Sorensen, NPA 855 (2011) 241;

B. Alver and G. Roland, PRC 81 (2010) 054904;

Planck data: EAS and the Planck collaboration

QGP plot: B. Schenke, S. Jeon, and C. Gale, arXiv:1109.6289.



Lesson II

- Geometrical anisotropy is seed for elliptic flow v_2
- Elliptic flow v_2 sensitive to QGP equation of state
- Triangular flow v_3 due to fluctuations, e.g. in initial energy density
- Triangular flow v_3 especially sensitive to shear viscosity /entropy density ratio
- Higher harmonics v_4, v_5, \dots strongly damped



Hydrodynamical model description

Some basic concepts

Relativistic Hydrodynamics (I)

The energy-momentum tensor $T^{\mu\nu}$ is the four-momentum component in the μ direction per three-dimensional surface area perpendicular to the ν direction.

$$\Delta \mathbf{p} = (\Delta E, \Delta p_x, \Delta p_y, \Delta p_z)$$

$$\Delta \mathbf{x} = (\Delta t, \Delta x, \Delta y, \Delta z)$$

$$\mu = \nu = 0 : T_R^{00} = \frac{\Delta E}{\Delta x \Delta y \Delta z} = \frac{\Delta E}{\Delta V} = \varepsilon$$

$$\mu = \nu = 1 : T_R^{11} = \frac{\Delta p_x}{\Delta t \Delta y \Delta z}$$

force in x direction acting on an surface $\Delta y \Delta z$ perpendicular to the force \rightarrow pressure

$$T^{\mu\nu} = \begin{pmatrix} \text{energy density} & \text{energy flux density} \\ \text{momentum density} & \text{momentum flux density} \end{pmatrix} \equiv \begin{pmatrix} \varepsilon & \vec{j}_\varepsilon \\ \vec{g} & \Pi \end{pmatrix}$$

Relativistic Hydrodynamics (II)

Isotropy in the fluid rest implies that the energy flux T^{0j} and the momentum density T^{j0} vanish and that $\Pi^{ij} = P \delta_{ij}$

$$T_R^{\mu\nu} = \begin{pmatrix} \varepsilon & 0 & 0 & 0 \\ 0 & P & 0 & 0 \\ 0 & 0 & P & 0 \\ 0 & 0 & 0 & P \end{pmatrix}$$

Off-diagonal elements $\neq 0$ in case of viscous hydrodynamics, not considered here
→ ideal (perfect) fluid.

See also Ollitrault, arXiv:0708.2433.

Relativistic Hydrodynamics (III)

Energy-momentum tensor (in case of local thermalization) after Lorentz transformation to the lab frame:

$$T^{\mu\nu} = (\varepsilon + P) u^\mu u^\nu - P g^{\mu\nu} \quad \text{metric tensor diag}(1, -1, -1, -1)$$

Energy density and pressure in the co-moving system

4-velocity: $u^\mu = dx^\mu / d\tau = \gamma(1, \vec{v})$

Energy and momentum conservation:

$$\partial_\mu T^{\mu\nu} = 0, \quad \nu = 0, \dots, 3$$

in components: $\left\{ \begin{array}{l} \frac{\partial}{\partial t} \varepsilon + \vec{\nabla} \cdot \vec{j}_\varepsilon = 0 \quad (\text{energy conservation}) \\ \frac{\partial}{\partial t} g_i + \nabla_j \Pi_{ij} = 0 \quad (\text{momentum conservation}) \end{array} \right.$

$\partial_\mu = \left(\frac{\partial}{\partial t}, \vec{\nabla} \right)$

Conserved quantities, e.g., baryon number:

$$j_B^\mu(x) = n_B(x) u^\mu(x), \quad \partial_\mu j_B^\mu(x) = 0 \Leftrightarrow \frac{\partial}{\partial t} N_B + \vec{\nabla} \cdot (N_B \vec{v}) = 0$$

continuity equation

$$N_B = \gamma n_B$$

Ingredients of Hydro - models

- Equation of motion and baryon number conservation:

$$\partial_\mu T^{\mu\nu} = 0, \quad \partial_\mu j_B^\mu(x) = 0$$

- 5 equations for 6 unknowns:

$$(u_x, u_y, u_z, \varepsilon, P, n_B)$$

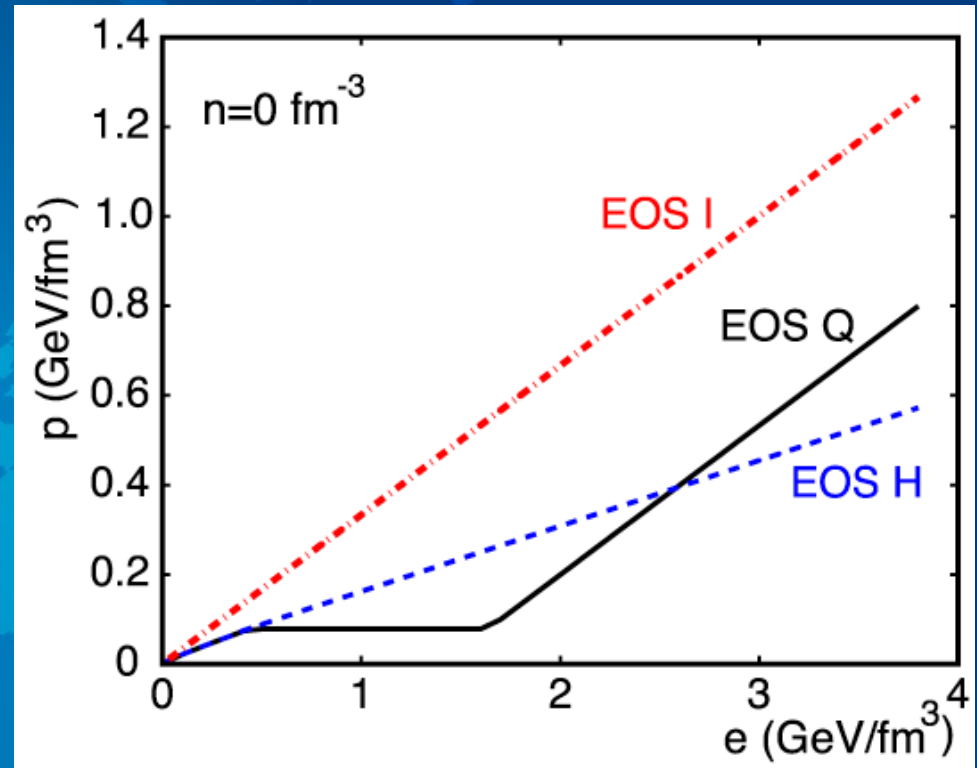
- Equation of state: $P(\varepsilon, n_B)$

- (needed to close the system)

- Initial conditions,

e.g., from Glauber calculation

- Freeze-out condition



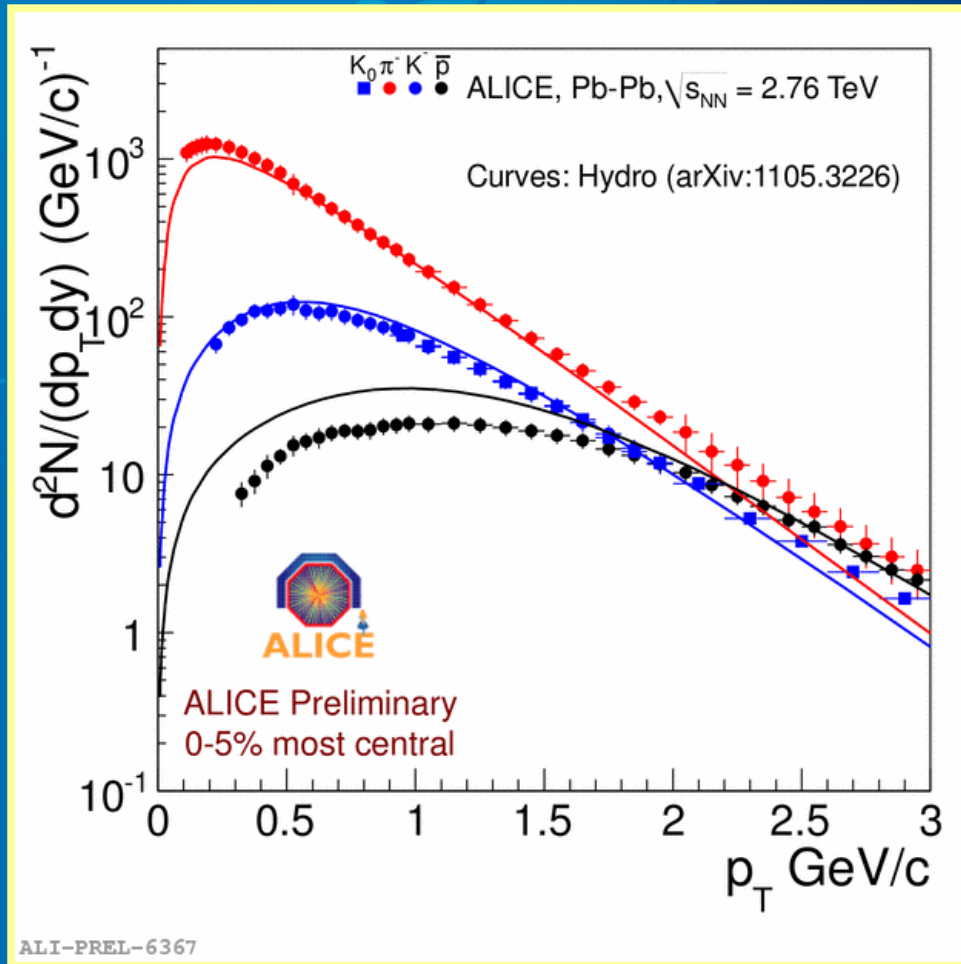
EOS I: ultra-relativistic gas $P = \varepsilon/3$

EOS H: resonance gas, $P \approx 0.15 \varepsilon$

EOS Q: phase transition,

QGP resonance gas

LHC: Identified particle spectra



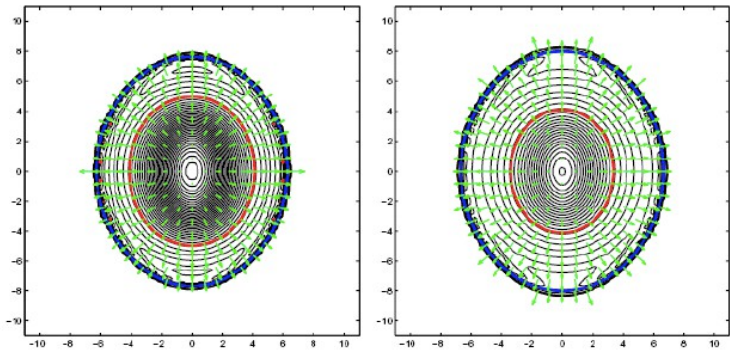
Initial conditions fixed
by pion abundance

Protons overestimated

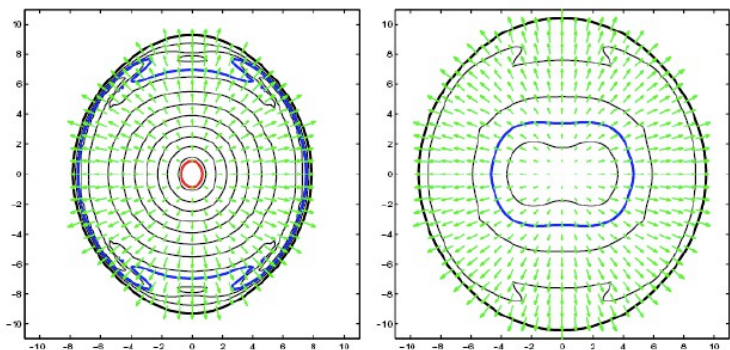
Annihilation of protons
and anti-protons in the
hadron phase ?

Elliptic flow in Hydro - models

Au+Au at $b = 7$ fm



3.2 fm/c ($\epsilon_x = 0.160, \epsilon_p = 0.114$) 4.0 fm/c ($\epsilon_x = 0.127, \epsilon_p = 0.141$)

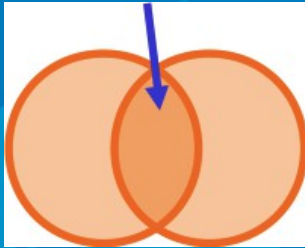


5.6 fm/c ($\epsilon_x = 0.067, \epsilon_p = 0.147$) 8.0 fm/c ($\epsilon_x = 0.003, \epsilon_p = 0.123$)

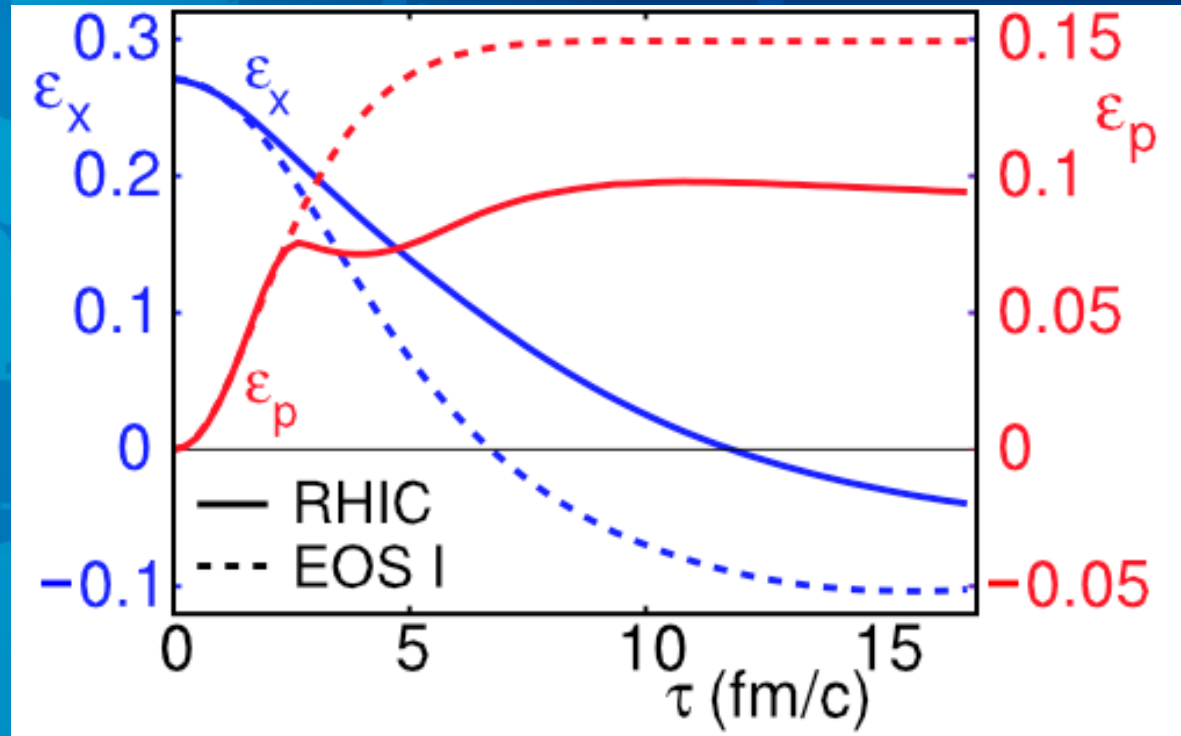
Elliptic flow is “self-quenching”: The cause of elliptic flow, the initial spacial anisotropy, decreases as the momentum anisotropy increases

Anisotropy in momentum space

Anisotropy in coordinate space



Anisotropy in momentum space



Ulrich Heinz, Peter Kolb, arXiv:nucl-th/0305084

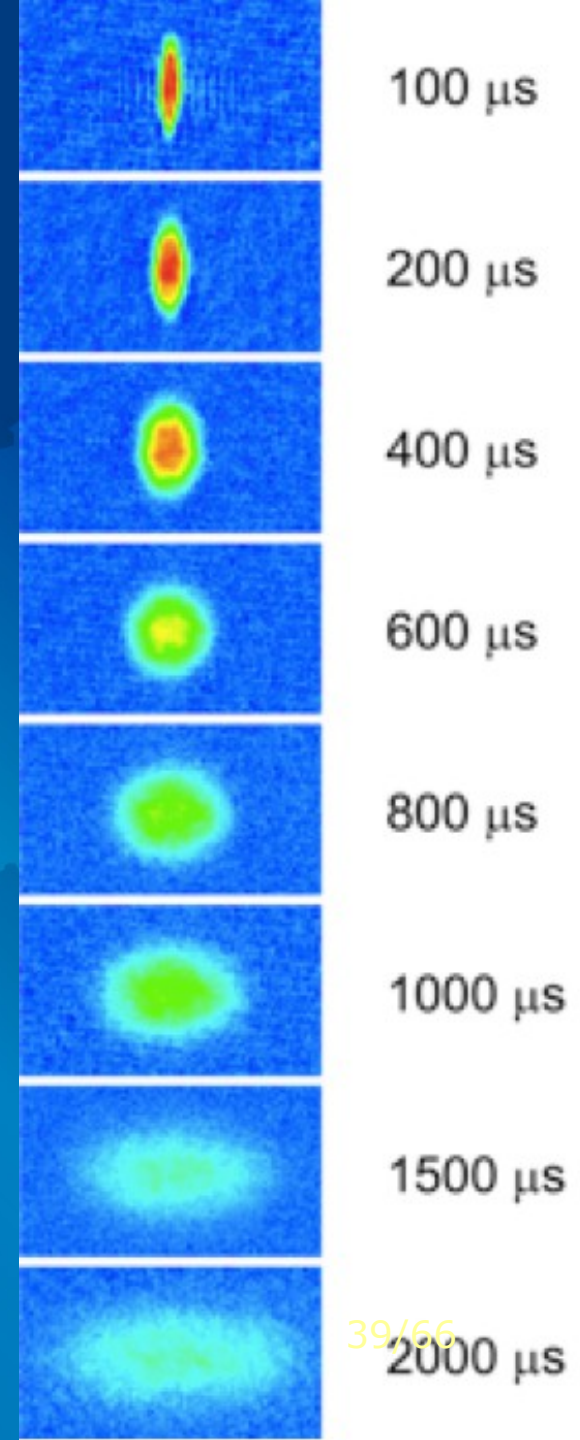
In hydrodynamic models the momentum anisotropy develops in the early (QGP) phase of the collision. Thermalization times of less than 1 fm/c are needed to describe the data.

Cold atomic gases

200 000 Li-6 atoms in an highly anisotropic trap (aspect ratio 29:1)

Very strong interactions between atoms (Feshbach resonance)

Once the atoms are released the one observed a flow pattern similar to elliptic flow in heavy-ion collisions



Lesson III

- **First results** from **ALICE** show large **increase** in **energy density** (**factor 2-3** compared to RHIC)
- **longer life-time** of qgp
- **larger collective flow** effects
- **anisotropic flow** comparable to **ultra-low viscosity**
- triangular flow sensitive to initial energy density fluctuations and viscosity/entropy ratio
- Hydrodynamical model provides framework to characterize QGP, i.e. equation of state, viscosity/entropy ratio